

3D evolution of a solar flare thermal X-ray loop-top source

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The Sexy Part



The Sexy Part

What?
How?
Why?

How Do We Reconstruct Sources in 3-D?
Why Is This New?

Magician's Mistake: Secrets Revealed

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Obs. 1

Obs.2

X-ray source

Magician's Mistake: Secrets Revealed



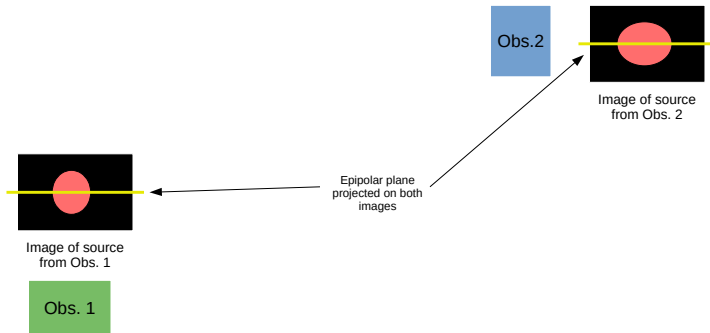
Image of source
from Obs. 1

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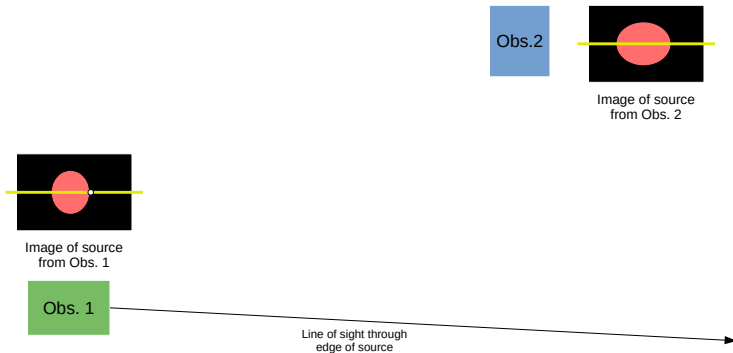


Image of source
from Obs. 2

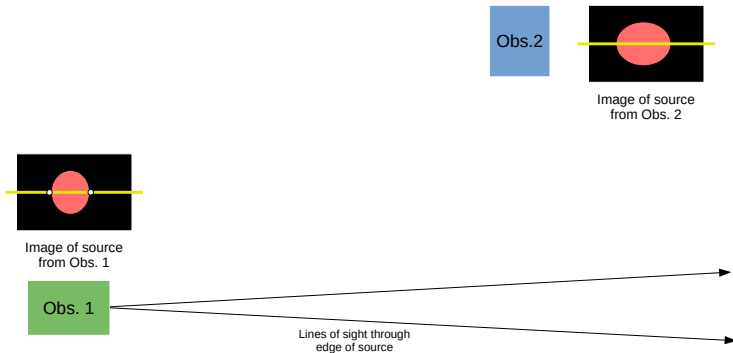
Magician's Mistake: Secrets Revealed



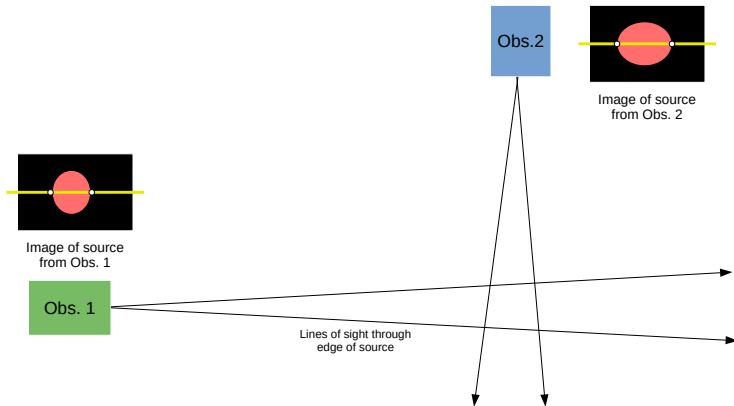
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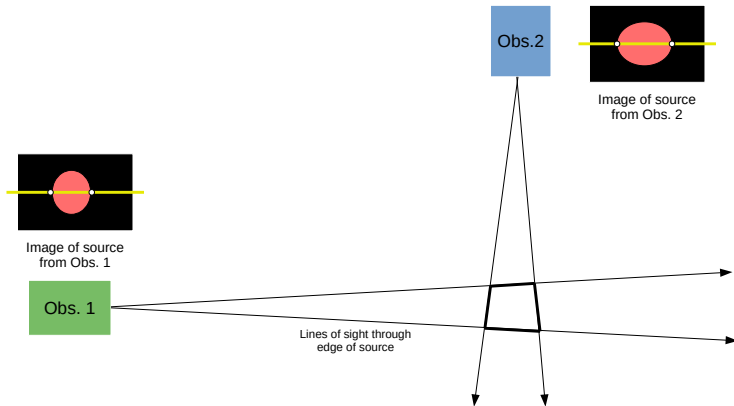
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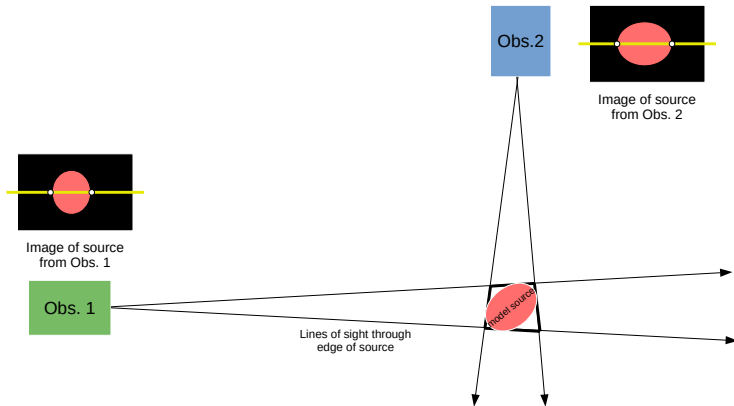
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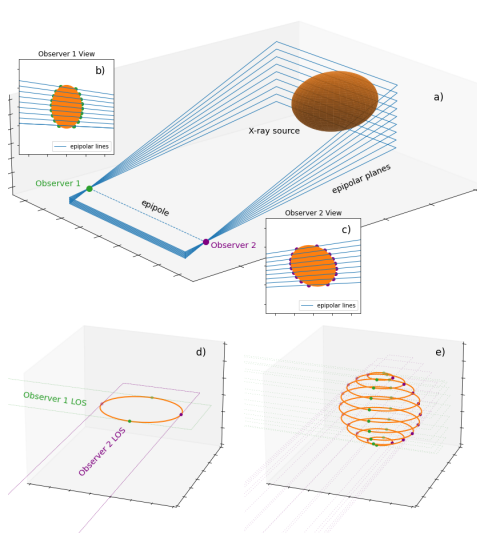
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Why Is This New?

What Do We Need?

Two X-ray telescopes with:

- Substantially different viewing angles;
- Same/similar passbands.

This has not previously been available...

Why Is This New?

What Do We Need?

Two X-ray telescopes with:

- Substantially different viewing angles;
- Same/similar passbands.

This has not previously been available...

...until recently!

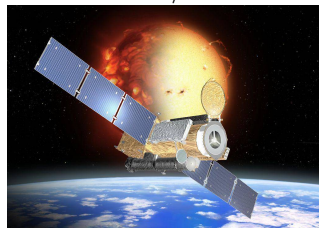
- Solar Orbiter/STIX 6–10 keV ($\gtrsim 8$ MK)
- Hinode/XRT Be-thick filter ($\gtrsim 5$ MK)

Comparing STIX & XRT for X-ray 3-D Reconstruction

Solar Orbiter/STIX



Hinode/XRT



	Solar Orbiter/STIX	Hinode/XRT
Viewing Angle	Variable	Earth
Passbands	Spectral Imager (>4 keV)	Imaging filters
Temp. Sensitivity	≈ 8 MK	≈ 5 MK (Be-thick filter)
Angular Resolution	7"	2"
Spatial Resolution	1400 km (0.3 AU)	1420 km (1 AU)
Max. Cadence	0.5 s (intensity-dependent)	2 s

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- 1: 3-D Evolution of a Thermal Loop-top Source
- 2: How Does Geometry Impact Thermodynamic Evolution?
- 3: How Do Area-to-Volume Scaling Laws Perform?

Key Questions

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- 1 What is the 3-D evolution of a flare's thermal X-ray loop-top source?

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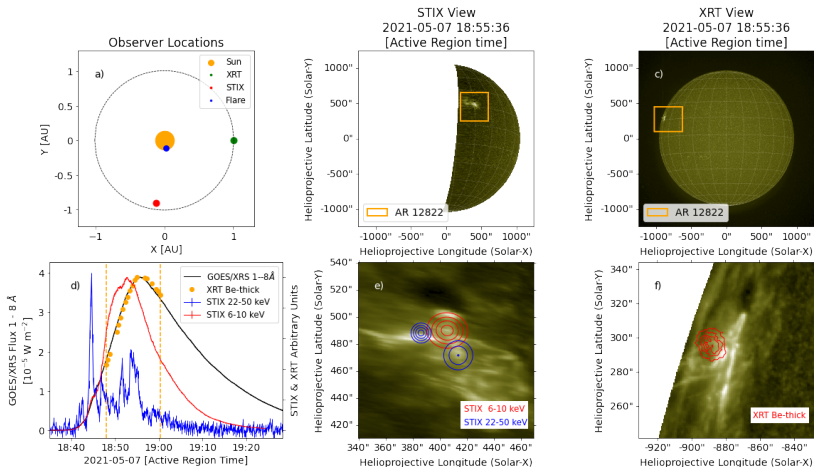
Key Questions

- 1 What is the 3-D evolution of a flare's thermal X-ray loop-top source?
- 2 How does the source's height and volume impact and its thermodynamic evolution?
- 3 How well do traditional area-to-volume scaling laws ($V \sim A^{3/2}$) approximate the 3-D volume?

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Result 1: 3-D Evolution of a Thermal Loop-top Source

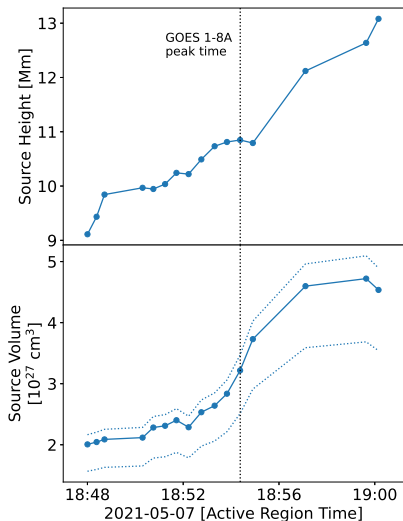


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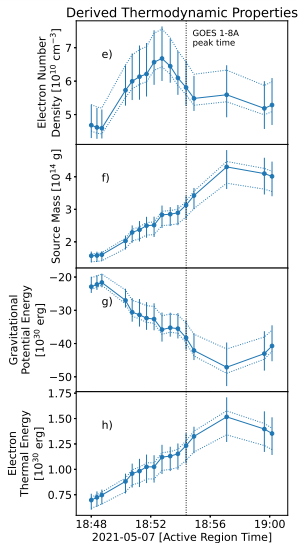
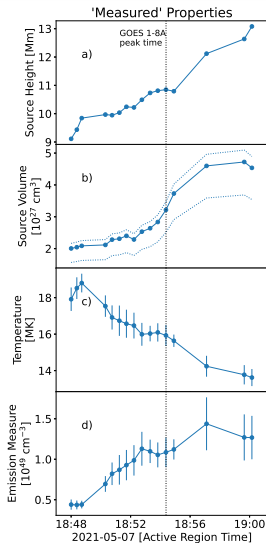
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Result 2: X-ray Source Thermodynamic Evolution



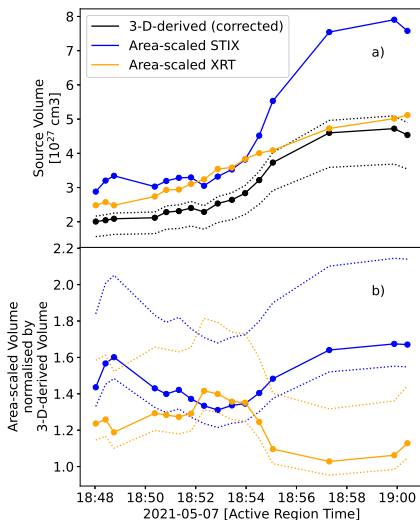
Result 3: How Do Area-to-volume Scaling Laws Perform?

Area-to-Volume Scaling Law

$$V = A^{3/2}$$

V: 3-D source volume

A: projected source area in image



Conclusions

For more, see Ryan et al. (2023, submitted)

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Conclusions

- 1 **STIX & XRT** enable 3-D reconstruction of flare thermal X-ray sources for the first time.
- 2 Area-to-volume scaling can overestimate the volume by up to a factor of 2, and **3-D analysis is required to capture asymmetric geometry evolution.**
- 3 **3-D analysis provides a way to quantify volume uncertainties.**
- 4 3-D analysis helps us **better understand the geometry, thermodynamics, energy transport, etc.** in solar flares, especially in multi-wavelength/model comparison studies.

For more, see Ryan et al. (2023, submitted)

Conclusions

Thank you for your attention!

For more, see Ryan et al. (2023)

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BACKUP SLIDES

IMPROVING VOLUME ESTIMATES AND DERIVING VOLUME UNCERTAINTIES

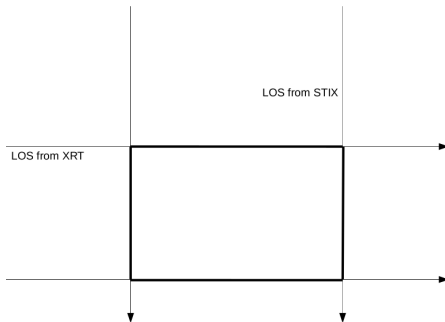
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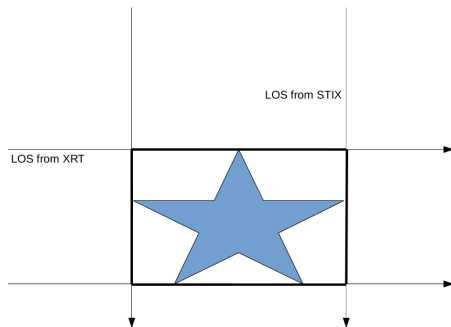
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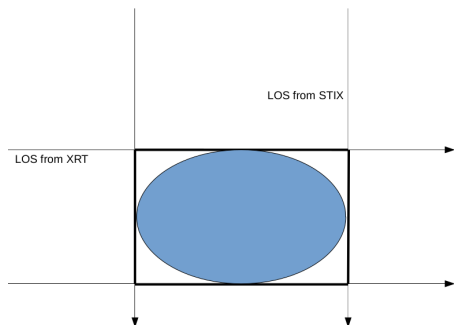


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Assumptions

- 1 The source cross-section is an ellipse.

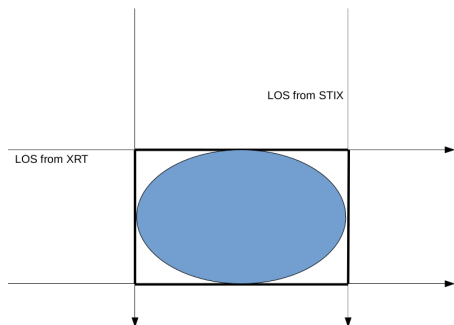


How Can We Improve Our 3-D Volume Estimates?

Simply integrating the between the reconstructed cross-sections overestimates the true volume due to inherent assumptions.

Assumptions

- 1 The source cross-section is an ellipse.
- 2 The ellipse occupies the maximum possible area within the bounding box.



How Can We Improve Our 3-D Volume Estimates?

Resulting Caveats

- Derived geometry is an approximation.
- Cross-sectional area are upper limits.

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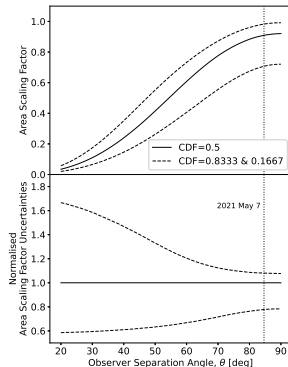
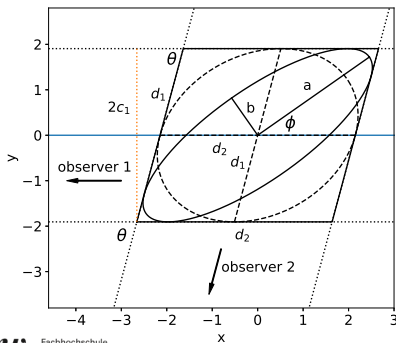
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However, we can improve the volume estimates.

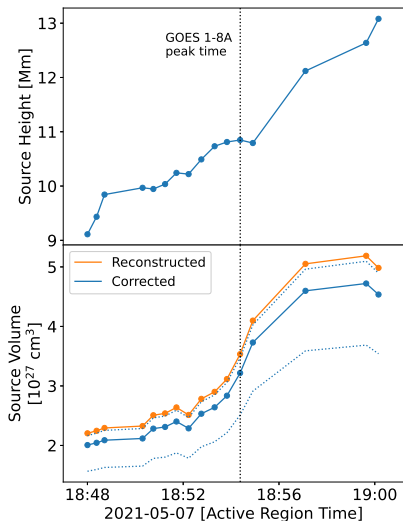
How Can We Improve Our 3-D Volume Estimates?

$$A_0 = \kappa A'; \quad \kappa = \frac{\sin^2 \theta}{\sqrt{(\sin^2 \phi + \rho^2 \cos^2 \phi) \left[\frac{1}{\rho^2} \left(\frac{\sin \phi}{\tan \theta} - \cos \phi \right)^2 + \left(\frac{\cos \phi}{\tan \theta} + \sin \phi \right)^2 \right]}}$$

A_0 : True cross-sectional area; A' : Derived cross-sectional area; $\rho = b/a$



Result 1: 3-D Evolution of a Thermal Loop-top Source



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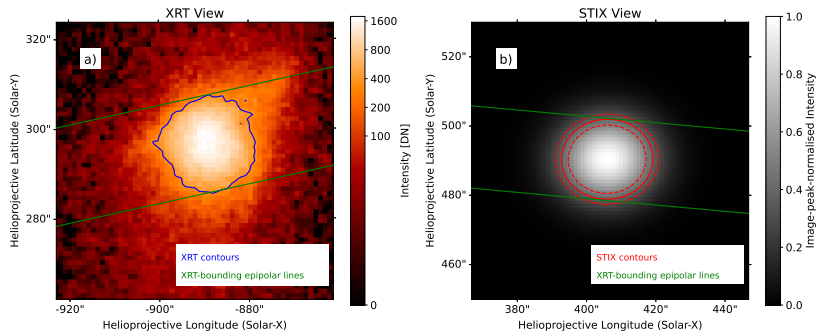
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DEFINING SOURCE BOUNDARIES

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How To Define Consistent Source Boundaries in STIX & XRT Images



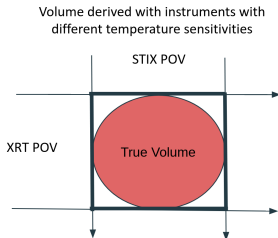
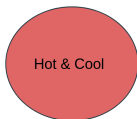
DO STIX & XRT SEE THE SAME VOLUME?

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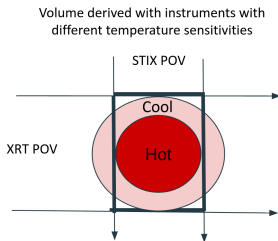
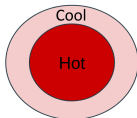
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Do STIX & XRT See the Same Volume?

Spatially homogeneous
temperature distribution



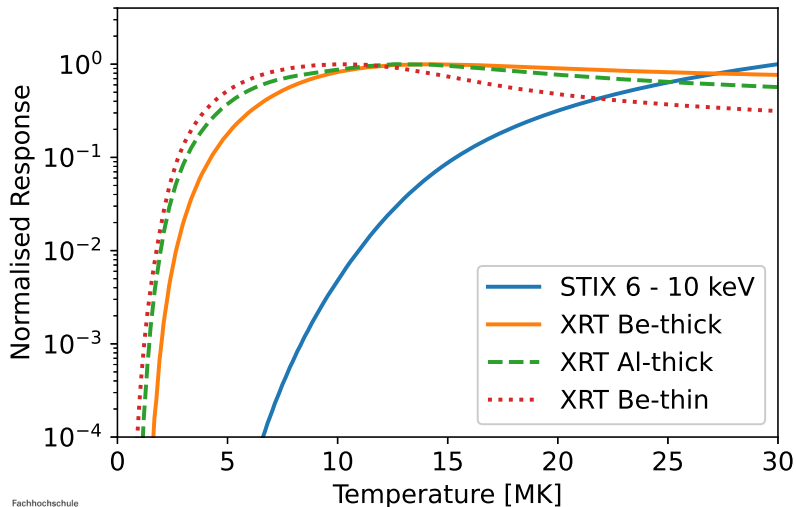
Spatially structured
temperature distribution



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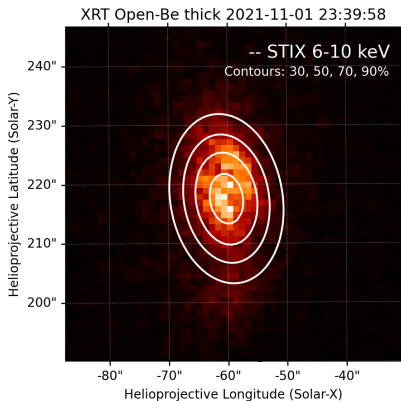
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STIX & XRT Sources Areas From Same Viewing Angle?



In at least some cases, XRT & STIX see the same flare volume.

Can STIX T & EM Predict the XRT Intensity?

$$I_{flare} = R(T) \frac{EM_V}{A_{pix}}$$

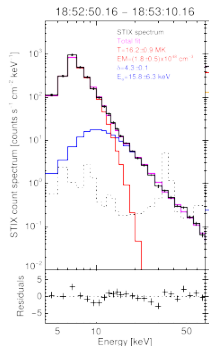
I_{flare} = Flare-summed XRT intensity

R = XRT temperature response

A_{pix} = area of XRT pixel at source

T = flare temperature

EM_V = flare volume emission measure



$$I_{xrt}^{pred} = 4500 \text{ DN/s}$$

$$I_{xrt}^{obs} = 3700 \text{ DN/s}$$

Agreement within ~20%

The plasma seen by STIX is consistent with the XRT observations.